

Commercializing a Next-Generation Source of Safe Nuclear Energy

Low Energy Nuclear Reactions (LENRs)

**Nucleosynthetic Networks Beginning with Nickel ‘Seed’ Nuclei,
Why Cascades of Fast Beta-Decays are Important, and
Why end-products of LENR networks are mostly stable isotopes**

Technical Overview

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***“Out of intense complexities ,
intense simplicities emerge.”***

Winston Churchill



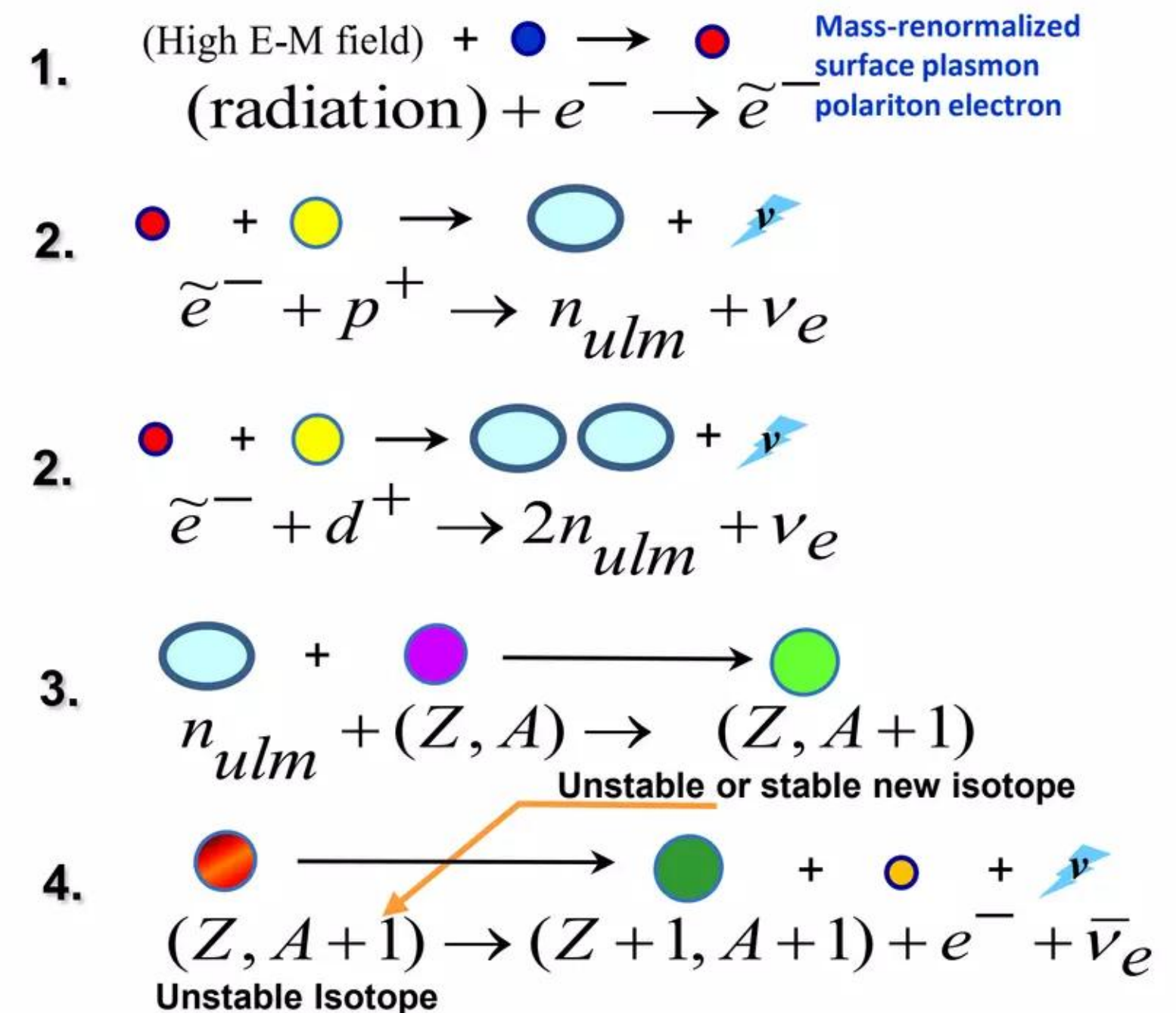
W-L mechanism in condensed matter LENR systems

Weak interaction processes are very important in LENRs

1. E-M radiation on metallic hydride surface increases mass of surface plasmon electrons
2. Heavy-mass surface plasmon polariton electrons react directly with surface protons (p^+) or deuterons (d^+) to produce ultra low momentum (ULM) neutrons (n_{ulm} or $2n_{ulm}$, respectively) and an electron neutrino (ν_e)
3. Ultra low momentum neutrons (n_{ulm}) are captured by nearby atomic nuclei (Z, A) representing some element with charge (Z) and atomic mass (A). ULM neutron absorption produces a heavier-mass isotope ($Z, A+1$) via transmutation. This new isotope ($Z, A+1$) may itself be a stable or unstable, which will perform eventually decay
4. Many unstable isotopes β^- decay, producing: transmuted element with increased charge ($Z+1$), \sim same mass ($A+1$) as 'parent' nucleus; β^- particle (e^-); and an antineutrino

→ Note: colored shapes associated with diagram on next Slide

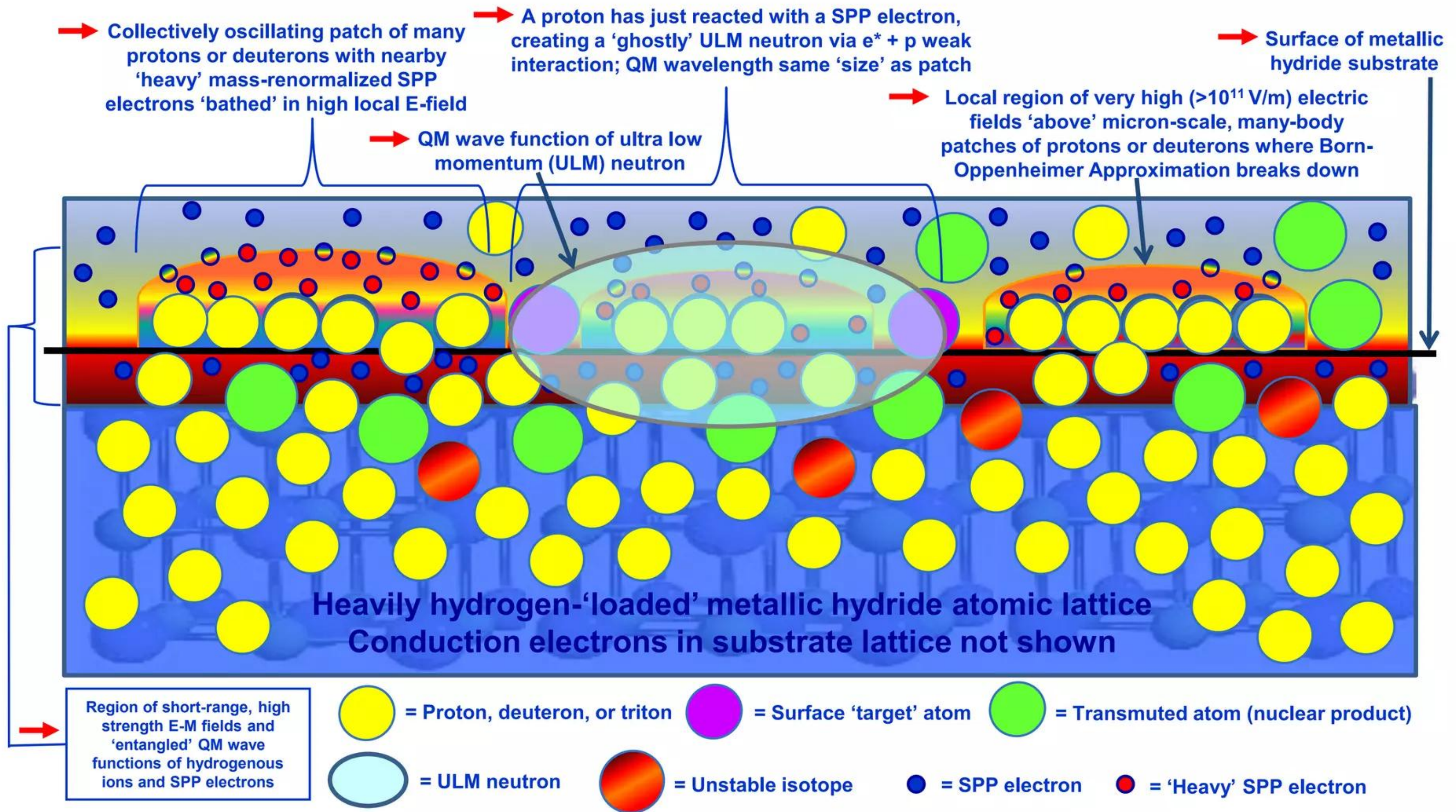
→ No strong interaction fusion or heavy element fission occurring below; weak interaction $e + p$ or $e + d$



→ Weak interaction β^- decays (shown above), direct gamma conversion to infrared (not shown), and α decays (not shown) produce most of the excess heat calorimetrically observed in LENR systems

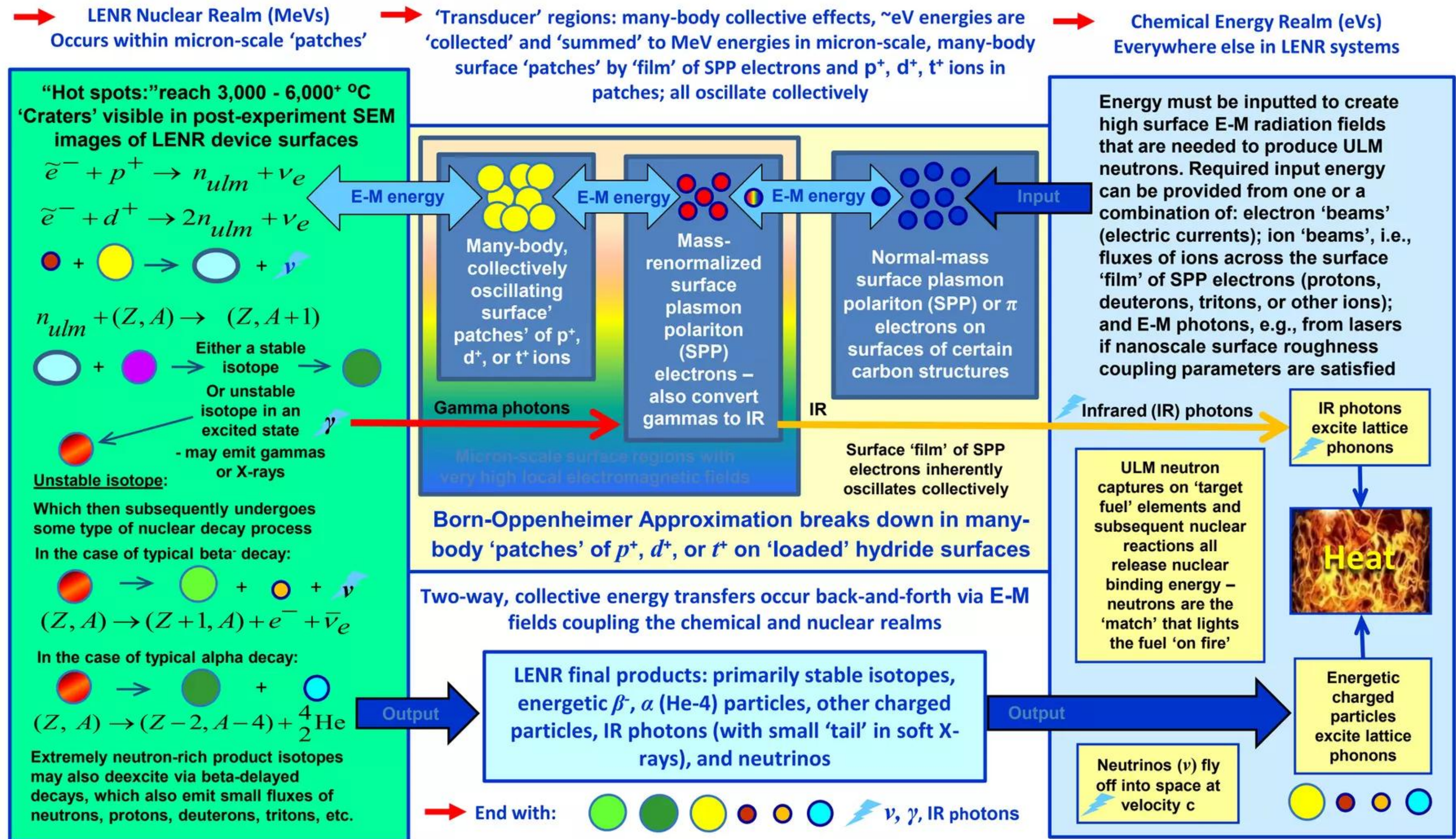
Conceptual details: W-L mechanism in metallic hydrides

Side view – not to scale – charge balances in diagram only approximate



High level overview: W-L mechanism in condensed matter

Chemical and nuclear energy realms can interconnect in small regions



Technical side note re ULM neutron capture cross-sections

- ✓ Unlike energetic neutrons produced in most nuclear reactions, collectively produced LENR neutrons are effectively 'standing still' at the moment of their creation in condensed matter. Since they are vastly below thermal energies (ultra low momentum), ULM neutrons have huge DeBroglie wavelengths (from *nm* to *~100 microns*) and accordingly large capture cross-sections on nearby nuclei; most or all will be locally absorbed; few will be detectable as 'free' neutrons
- ✓ For the vast majority of stable and unstable isotopes, their neutron capture cross-section (relative to measurements of cross-sections at thermal energies where $v = 2,200 \text{ m/sec}$ and neutron DeBroglie wavelength is *~2 Angstroms*) is proportional to $\sim 1/v$, where v is velocity of a neutron in *m/sec*. Since v is extraordinarily small for ULM neutrons, their capture cross-sections on atomic nuclei will therefore be correspondingly larger. After being collectively created, an enormous percentage of the ULMNs produced will be locally absorbed before scattering on nearby atoms can elevate them to thermal kinetic energies; per Prof. S. Lamoreaux (Yale) thermalization would require *~0.1 to 0.2 msec*, i.e. 10^{-4} sec. , a very long time on typical $10^{-16} - 10^{-19} \text{ sec.}$ time-scale of nuclear reactions

*Please note: ultra low momentum (ULM) neutrons have enormous absorption cross-sections on $1/v$ isotopes. For example, Lattice has estimated the ULMN fission capture cross-section on U-235 to be *~1 million barns (b)* and on Pu-239 at *49,000 b*, vs. *~586 b* and *~752 b*, respectively, for 'typical' neutrons at thermal energies*

*A neutron capture expert recently estimated the ULMN capture cross-section on He-4 at *~20,000 b* vs. a value of *<1 b* for thermal neutrons; this is a huge increase*

*By comparison, the highest known thermal n capture cross section for any stable isotope is Gadolinium-157 at *~49,000 b**

*The highest measured cross-section for any unstable isotope is Xenon-135 at *~2.7 million b**

Crucial technical point: ULMNs have many-body scattering, NOT 2-3 body scattering as, for example, in plasmas or thermalized neutrons traveling through condensed matter

Decays of Neutron-rich Halo Nuclei More Complicated - I

Can dynamically vary their decay 'choices' depending on their 'environment'





- ✓ “Excited states of nuclei formed in beta decay, for example, can show other types of radioactive behavior. In such beta-delayed radioactivity, the excited nuclear states can emit other particles ... If, however, the excitation [of the ‘daughter’ nucleus] is high enough, then it is possible that an alpha particle, a neutron or a proton are emitted from this state... This radioactivity is beta-delayed because the particles are only emitted after a time equal to the half-life of the beta particle [emission].” W. Scharf, “Particle Accelerators and Their Uses,” pp. 726 Taylor & Francis 1991.
- ✓ Few observations of beta-delayed particles published prior to 1965
- ✓ 12 types of beta-delayed particle emissions have been observed
- ✓ Over 100 beta-delayed particle radioactivities are known today
- ✓ Theoretically, perhaps >1,000 isotopes could exhibit such decays

Decays of Neutron-rich Halo Nuclei More Complicated - II

Can dynamically vary their decay 'choices' depending on their 'environment'

Per W-L theory, once ULM neutron production begins at high rates, populations of unstable, very neutron-rich 'halo' isotopes build-up locally on 2-D surfaces. Such nuclei likely have substantially lengthened half-lives because they may have a difficult time emitting beta electrons or neutrons (both of which are fermions) into locally unoccupied Q-M states. By contrast, alpha (He-4) particle and gamma photon emissions are bosons and are unaffected by the exclusion principle. If fermionic decay channels are 'blocked', neutron-rich halo nuclei may emit bosons or continue capturing ULM neutrons as long as it is energetically favorable or until they finally get so neutron-rich and excited, or a previously occupied local state opens-up, that 'something breaks' and β^- decay cascades ending in stable isotopes can begin. This is one important reason why LENR systems typically do not end-up with large amounts of long-lived, radiologically 'hot' isotopes

Importantly, the neutron-capture phase of LENRs can release substantial amounts of nuclear binding energy, much of it in the form of prompt and delayed gammas (which are bosons). Unique to LENR systems and according to W-L theory, those gammas are converted directly to infrared photons by heavy SP electrons also present in nuclear-active 'patches' on surfaces in LENR systems. As explained elsewhere, beta-decay cascades of unstable isotopes with short half-lives can proceed very rapidly, release large amounts of binding energy, and produce complex arrays of different transmutation products that, if neutron fluxes are high enough, can rapidly traverse rows of the periodic table; in one spectacular experiment, Mizuno went from K to Fe in <2 minutes

Weak interaction	W-L neutron production	LENR Nuclear Realm (MeVs) Occurs within micron-scale 'patches' $\tilde{e}^- + p^+ \rightarrow n_{ulm} + \nu_e$ $\tilde{e}^- + d^+ \rightarrow 2n_{ulm} + \nu_e$ 
Strong interaction	Neutron capture	$n_{ulm} + (Z, A) \rightarrow (Z, A+1)$  Either a: stable or unstable HEAVIER isotope
Transmutations: isotope shifts occur; chemical elements disappear/appear	Decays of unstable, very neutron-rich isotopes: beta and alpha (He-4) decays	<u>In the case of unstable isotopic products:</u> they subsequently undergo some type of nuclear decay process; e.g., beta, alpha, etc. In the case of a typical beta ⁻ decay:  $(Z, A) \rightarrow (Z+1, A) + e^- + \bar{\nu}_e$ In the case of a typical alpha decay:  $(Z, A) \rightarrow (Z-2, A-4) + {}^4_2\text{He}$ <u>Note:</u> extremely neutron-rich product isotopes may also deexcite via beta-delayed decays, which can also emit small fluxes of neutrons, protons, deuterons, tritons, etc.

Decays of Neutron-rich Halo Nuclei More Complicated - III

Can dynamically vary their decay 'choices' depending on their 'environment'

- ✓ At this point in our understanding of nuclear physics, note that extremely neutron-rich 'halo' nuclei that comprise intermediate products created in condensed matter LENR nucleosynthetic networks are, in many respects, still little understood and poorly characterized. For example: neutron capture cross-sections are unknown for many; short half-lives can be difficult to measure accurately; true location of neutron 'dripline' unclear at higher A
- ✓ In condensed matter LENRs, neutron-rich halo isotopes may continue to absorb ULM neutrons as long as capture Q-values remain favorable (prompt and delayed capture gammas are converted into infrared by heavy SP electrons) and as long as they are unable to decay via a variety of available channels that include emitting neutrinos and β^- electrons (fermions) and/or shedding low-energy neutrons (fermions) into unoccupied states in local continuum. Importantly, *LENR systems have much higher local densities of occupied fermionic states than various neutron-rich fragments made in radioactive-ion beam collider experiments*

Decays of Neutron-rich Halo Nuclei More Complicated - IV

Can dynamically vary their decay 'choices' depending on their 'environment'

- ✓ Key consequences of this unique situation in LENR systems (very dense occupation of local fermionic states) are that: (1) effective half-lives of very neutron-rich intermediate isotopes can sometimes be substantially longer than measured 'textbook' half-lives of comparatively isolated excited nuclei; and (2) % branching ratios for alternative β -delayed decay channels that are normally 'available' to such isotopes may change markedly compared to isolated nuclei – branching ratios can shift dynamically if decay channels are 'blocked,' i.e., unavailable. Thus certain types of decays can be 'frustrated' in LENR systems until unoccupied states 'open-up,' for whatever reason
- ✓ From low (better understood) to high values of A, unstable neutron-rich isotopes far from the valley of stability have >> greater variety of decay channel 'choices' compared to less neutron-rich nuclei. These include emissions of *beta*⁻ electrons along with delayed gammas, neutrons (up to 3), protons, alpha particles (He-4), tritons (Tritium), and/or deuterons (Deuterium). Although production cross-sections often small, in some isotopes they are large, e.g., in N-18 ~12% of β -decays also emit He-4

Extensive evidence for transmutations via W-L mechanism

✓ Utilizing the W-L theory of LENRs, readers are encouraged to examine reported experimental data and evaluate evidence for transmutations and nuclear energy releases in LENR systems

✓ In doing so, please be aware that :

- Any element or isotope present in LENR experimental apparatus having an opportunity to move into close physical proximity to surfaces or nanoparticles on which ULM neutrons are being created can potentially ‘compete’ with other nuclei (that are located within the same micron-scale domains of spatially extended ULM neutron Q-M wave functions) to absorb produced ULMNs
- Some reported transmutation products may appear rather mystifying until one determines *exactly what elements/isotopes were initially present in the apparatus when an experiment began*. In many cases, materials located inside experiments are poorly characterized; thus ‘starting points’ for ULMN captures on initial ‘seed’ nuclei may be very unclear

There is an extensive body of experimental data on LENRs. Some of it can be found via the following resources:

- *References cited in publications by Widom, Larsen, and Srivastava on the W-L theory of LENRs*
- *Go to the free website www.lenr-canr.org Several hundred .pdf downloadable papers are available on site (**most documents are not peer-reviewed**)*
- *Go to the free website www.newenergytimes.com Investigative articles and news, as well as number of downloadable papers on site*
- *Query Internet search engines like Google using various key words such as: “low energy nuclear reactions”, “cold fusion energy -Adobe”, “cold fusion”, and so forth*

Extensive evidence for transmutations via W-L mechanism

Ca. 2002, Miley's informal survey of researchers' experiments reveals that Fe (#1), Zn (#2), and Cu (#3) most commonly reported LENR transmutation products

Frequency	Transmutation Elements
1	23 (F, Lu, Tb, I, Br, Xe, Os, Pr, Li, B, O, Sc, Ge, Se, Rb, Y, Zr, Eu, Sm, Gd, Dy, Ho, Nd)
2	14 (Ag, V, Yb, C, As, Sb, Te, Pd, Au, Cs, Mo, Ba, Nb, In)
3	9 (Cl, Hf, Re, Na, Ga, Sr, Sn, Cd, Ir)
5	4 (Co, S, Ti, Pt)
6	4 (Ni, K, Mn, Pb)
7	3 (Al, Mg, Si)
11	2 (Ca, Cr)
12	1 (Cu)
14	1 (Zn)
15	1 (Fe)

Source of Table: Circa 2002 by Prof. George Miley, Dept. of Nuclear Engineering,
University of Illinois at Urbana-Champaign

Extensive evidence for transmutations via W-L mechanism

Examples of results reported prior to 1st arXiv preprint of W-L theory (May 2005)

Isotopes	Miley, G.,	Ohmori, T. et al.	Ohmori, T. and Mizuno, M.	Mizuno, M and Enyo, M.,	Ohmori, T. et al.,	Chernov, I.P. et al.,	Iwamura, Y
	1997	2000	1998	1996	1998	1998	2002
Cu-63	3.6%±1.6			30.83%		44%	
Cu-65	-8.1%±3.6			-30.83%		-44%	
Fe-54		0%	-0.85%		-0.81%		
Fe-56		-0.75%	-29.75%	-21.75%	-16.79%		
Fe-57		0.78%	30.88%	18.88%	16.58%		
Fe-58					1.01%		
Re-185		11%		14.60%			
Re-187		-11%		-14.60%			
Ag-107	3.9%±1.2						
Ag-109	-4.3%±1.3						
Mo-96							26.26%
Ti-48						-33.30%	
Ti-49						8.60%	
Ti-50						17.80%	

Source: Circa 2003 by Prof. George Miley, Dept. of Nuclear Engineering, University of Illinois at Urbana-Champaign

Extensive evidence for transmutations via W-L mechanism

Ca. 2000, on the basis of experiments involving Nickel-cathode light water (Ni-H) P&F-type electrolytic cells, Miley reports claim of a rough correlation between calorimetrically measured excess heat and nuclear binding energy calculations

Figure 3.	Run Number (Ref. 2, 3)	Excess Power (W)	
		Calculated	Measured
	#7	1.9 ± 0.6	4.0 ± 0.8
	#8	0.5 ± 0.2	0.5 ± 0.4
	#18	0.7 ± 0.3	0.6 ± 0.4
Figure 3. Results from energy balance calculations for three earlier thin-film experiments. All experiments used Li_2SO_4 in H_2O for the electrolyte and thin-film Ni coated cathodes.			

Quoting discussion of Fig. 3 directly from the paper: “As seen from this figure, reasonably good agreement is obtained between the excess power measurement and the calculated values using the binding energy calculations described here. Two of the results show quite close agreement, but one has a mean measured value that is a factor of two larger than the calculated value. While these results are not definitive, still, in view of the many uncertainties in both of the calculated values (due to uncertainties in the yield measurements) and in the calorimetry, the agreement obtained strongly suggests a relation between products and excess heat.”

Source URL = <http://www.lenr-canr.org/acrobat/MileyGHonthereact.pdf>

**Miley, G.H., “On the Reaction Product and Heat Correlation for LENRs”
Presented at the 8th International Conference on Cold Fusion – 2000 in Lerici (La Spezia), Italy**

Extensive evidence for transmutations via W-L mechanism

Below is a facsimile of information copied from a slide presented by Prof. George Miley at a “Fusion Trends” conference ca. 2003

Summary - Key Experimental “Signatures” for LENRs



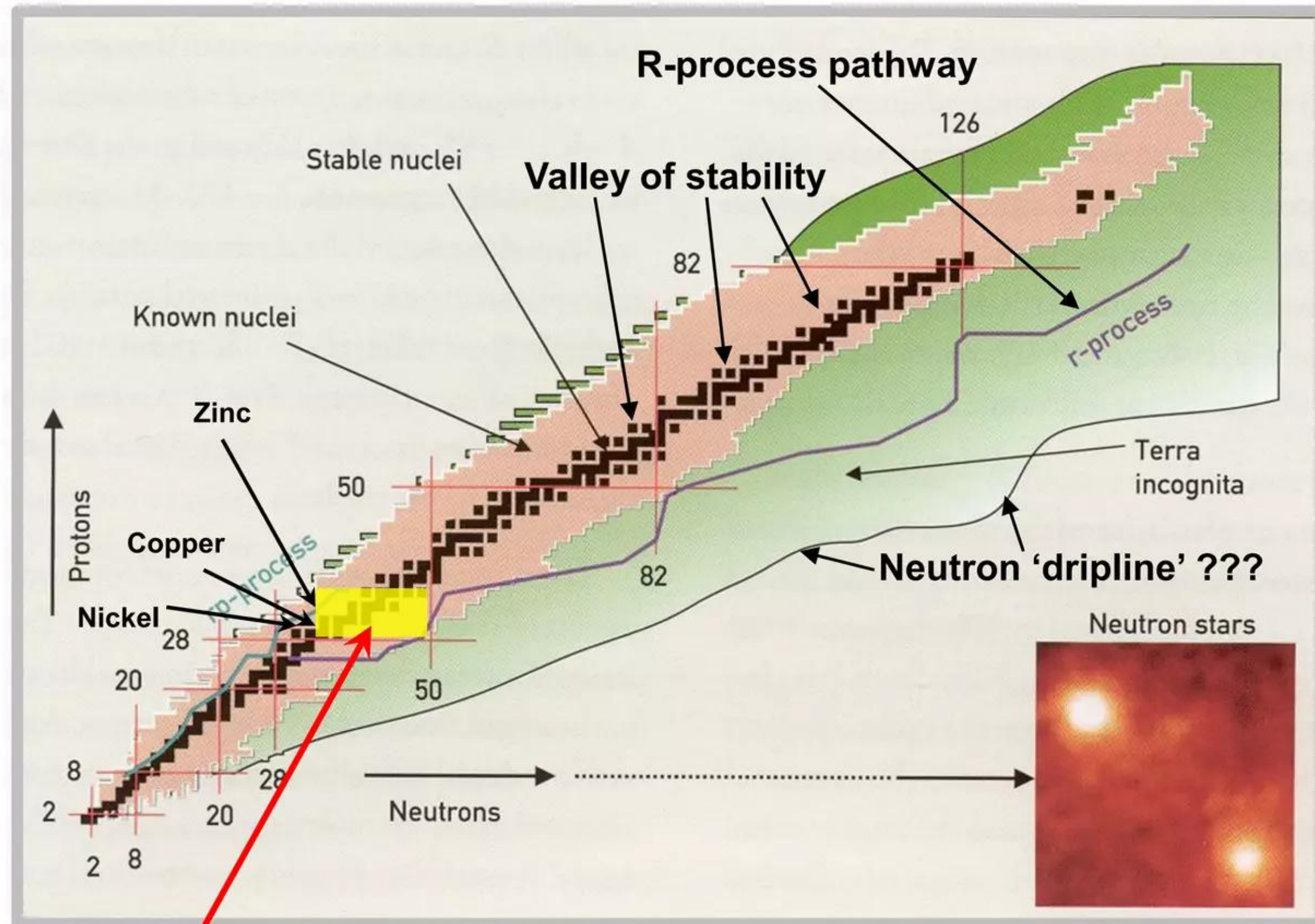
- Large reaction rates for key elements, e.g. Zn, Cu, Fe, Mg, and Ag.
- Peak mass zones of high yields separated by low yield “gaps”
- Non-natural isotope ratios for key products
- Lack of energetic n, gamma radiation, but
 - Low-energy (~20 keV) X-ray/beta radiation along with MeV level protons and alphas

Source of Slide: Prof. George Miley, Dept. of Nuclear Engineering,
University of Illinois at Urbana-Champaign

Preview of Nucleosynthetic Pathways - I

'Map' of Vast Isotopic Nuclear Landscape

The neutron-capture so-called "r-process" (see path on chart) that astrophysicists believe occurs mainly in stellar supernova explosions is thought to produce most of the nuclei heavier than Iron (Fe). It operates in the neutron-rich region of the nuclear landscape to the right of the valley of stability to beta⁻ decay. Extremely neutron-rich isotopes have a much wider variety of available decay channels in addition to 'simple' β^- .



While they differ from stellar environments in many important aspects, LENR systems can produce large fluxes of a wide variety of extremely neutron-rich nuclei from low to very high values of A. Thus, they may someday be able to provide nuclear physics with a new and exciting, much lower-cost experimental tool for exploring the far reaches of the nuclear landscape and boundaries of nuclear stability. This possibility deserves further careful study.

In this presentation, we will apply W-L theory and examine a model LENR Ni-seed nucleosynthetic network operating in the *approximate* yellow rectangular region from the valley of stability (small black squares) out thru very neutron-rich, beta⁻ decay isotopic regions that lie to the 'right' of valley of stability around Nickel (Ni), Copper (Cu), and Zinc (Zn)

Preview of Nucleosynthetic Pathways - II

THE PERIODIC TABLE

Major vector of LENR nucleosynthetic pathway discussed herein shown in red

Begin at Nickel (Ni)

H	SYMBOL
1	ATOMIC NUMBER
1.008	ATOMIC WEIGHT
Hydrogen	NAME

() = ESTIMATES

1 IA H 1 1.008 Hydrogen	2 IIA He 2 4.00 Helium																	18 VIIIA He 2 4.00 Helium
3 Li 3 6.94 Lithium	4 Be 4 9.01 Beryllium											13 IIIA B 5 10.81 Boron	14 IVA C 6 12.01 Carbon	15 VA N 7 14.01 Nitrogen	16 VIA O 8 16.00 Oxygen	17 VIIA F 9 19.00 Fluorine	18 VIIIA Ne 10 20.18 Neon	
11 Na 11 22.99 Sodium	12 Mg 12 24.31 Magnesium	3 IIIB	4 IVB	5 VB	6 VIB	7 VIIB	8 VIII	9 VIII	10 VIII	11 IB	12 IIB	13 IIIA Al 13 26.98 Aluminum	14 IVA Si 14 28.09 Silicon	15 VA P 15 30.97 Phosphorus	16 VIA S 16 32.07 Sulfur	17 VIIA Cl 17 35.45 Chlorine	18 VIIIA Ar 18 39.95 Argon	
19 K 19 39.10 Potassium	20 Ca 20 40.08 Calcium	21 Sc 21 44.96 Scandium	22 Ti 22 47.88 Titanium	23 V 23 50.94 Vanadium	24 Cr 24 52.00 Chromium	25 Mn 25 54.94 Manganese	26 Fe 26 55.85 Iron	27 Co 27 58.93 Cobalt	28 Ni 28 58.69 Nickel	29 Cu 29 63.55 Copper	30 Zn 30 65.39 Zinc	31 Ga 31 69.72 Gallium	32 Ge 32 72.61 Germanium	33 As 33 74.92 Arsenic	34 Se 34 78.96 Selenium	35 Br 35 79.90 Bromine	36 Kr 36 83.80 Krypton	
37 Rb 37 85.47 Rubidium	38 Sr 38 87.62 Strontium	39 Y 39 88.91 Yttrium	40 Zr 40 91.22 Zirconium	41 Nb 41 92.91 Niobium	42 Mo 42 95.94 Molybdenum	43 Tc 43 (97.9) Technetium	44 Ru 44 101.07 Ruthenium	45 Rh 45 102.91 Rhodium	46 Pd 46 106.42 Palladium	47 Ag 47 107.87 Silver	48 Cd 48 112.41 Cadmium	49 In 49 114.82 Indium	50 Sn 50 118.71 Tin	51 Sb 51 121.76 Antimony	52 Te 52 127.60 Tellurium	53 I 53 126.90 Iodine	54 Xe 54 131.29 Xenon	
55 Cs 55 132.91 Cesium	56 Ba 56 137.33 Barium	57 La 57 138.91 Lanthanum	72 Hf 72 178.49 Hafnium	73 Ta 73 180.95 Tantalum	74 W 74 183.85 Tungsten	75 Re 75 186.21 Rhenium	76 Os 76 190.2 Osmium	77 Ir 77 192.22 Iridium	78 Pt 78 195.08 Platinum	79 Au 79 196.97 Gold	80 Hg 80 200.59 Mercury	81 Tl 81 204.38 Thallium	82 Pb 82 207.2 Lead	83 Bi 83 208.98 Bismuth	84 Po 84 (209) Polonium	85 At 85 (210) Astatine	86 Rn 86 (222) Radon	
87 Fr 87 223.02 Francium	88 Ra 88 226.03 Radium	89 Ac 89 227.03 Actinium	104 Rf 104 (261) Rutherfordium	105 Db 105 (262) Dubnium	106 Sg 106 (263) Seaborgium	107 Bh 107 (262) Bohrium	108 Hs 108 (265) Hassium	109 Mt 109 (266) Meitnerium	110 Unnamed Discovery 110 Nov. 1994	111 Unnamed Discovery 111 Nov. 1994	112 Unnamed Discovery 112 1996	113 Unnamed Discovery 113 1999	114 Unnamed Discovery 114 1999	115 Unnamed Discovery 115 1999	116 Unnamed Discovery 116 1999	117 Unnamed Discovery 117 1999	118 Unnamed Discovery 118 1999	
ALKALI METALS	ALKALI EARTH METALS											HALOGENS					NOBLE GASES	

Mostly end-up at Copper (Cu) and Zinc (Zn)

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LANTHANIDES

58 Ce 140.12 Cerium	59 Pr 140.91 Praseodymium	60 Nd 144.24 Neodymium	61 Pm (145) Promethium	62 Sm 150.36 Samarium	63 Eu 152.97 Europium	64 Gd 157.25 Gadolinium	65 Tb 158.93 Terbium	66 Dy 162.50 Dysprosium	67 Ho 164.93 Holmium	68 Er 167.26 Erbium	69 Tm 168.93 Thulium	70 Yb 173.04 Ytterbium	71 Lu 174.97 Lutetium
90 Th 232.04 Thorium	91 Pa 231.04 Protactinium	92 U 238.03 Uranium	93 Np 237.05 Neptunium	94 Pu (240) Plutonium	95 Am 243.06 Americium	96 Cm (247) Curium	97 Bk (248) Berkelium	98 Cf (251) Californium	99 Es 252.08 Einsteinium	100 Fm 257.10 Fermium	101 Md (257) Mendelevium	102 No 259.10 Nobelium	103 Lr 262.11 Lawrencium

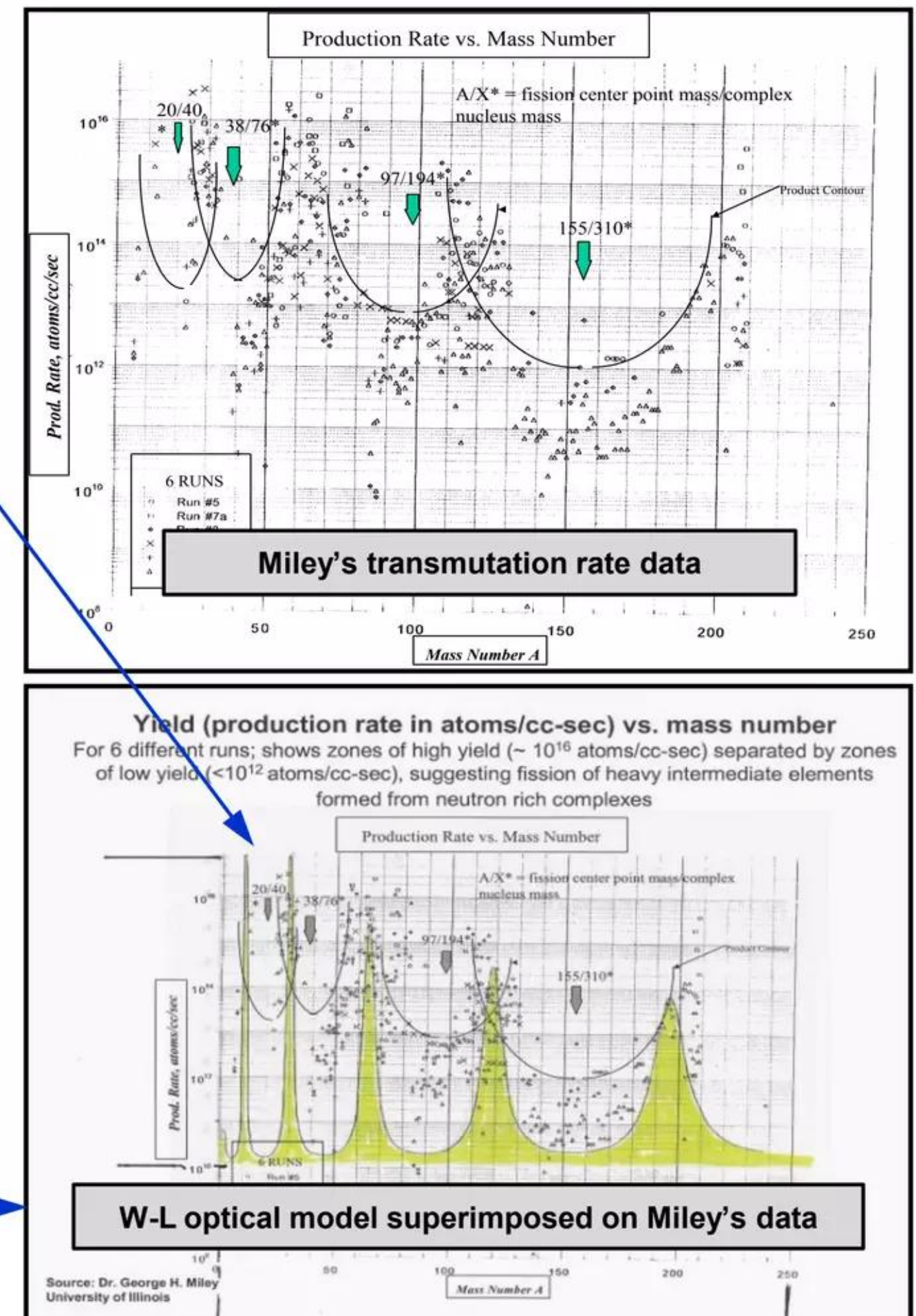
ACTINIDES

Five-peak mass-spectrum is a 'fingerprint' of ULM neutrons

Miley's dataset is 'smoking gun' for ULM neutron absorption by nuclei

- ✓ Top chart to right is Miley's raw data; chart below is same data only with results of W-L neutron optical potential model of ULMN neutron absorption by nuclei (yellow peaks) superimposed on top of Miley's data; good correspondence of Miley obs. vs. W-L calc.
- ✓ Model not fitted to data: only 'raw' calc output
- ✓ W-L model only generates a five-peak resonant absorption spectrum at the zero momentum limit; neutrons at higher energies will not produce the same result
- ✓ This means that 5-peak product spectrum experimentally observed by Miley and Mizuno is a unique 'signature' of W-L ULM neutron production and absorption (capture) in LENRs

See: "Nuclear abundances in metallic hydride electrodes of electrolytic chemical cells" arXiv:cond-mat/0602472 (Feb 2006) A. Widom and L. Larsen

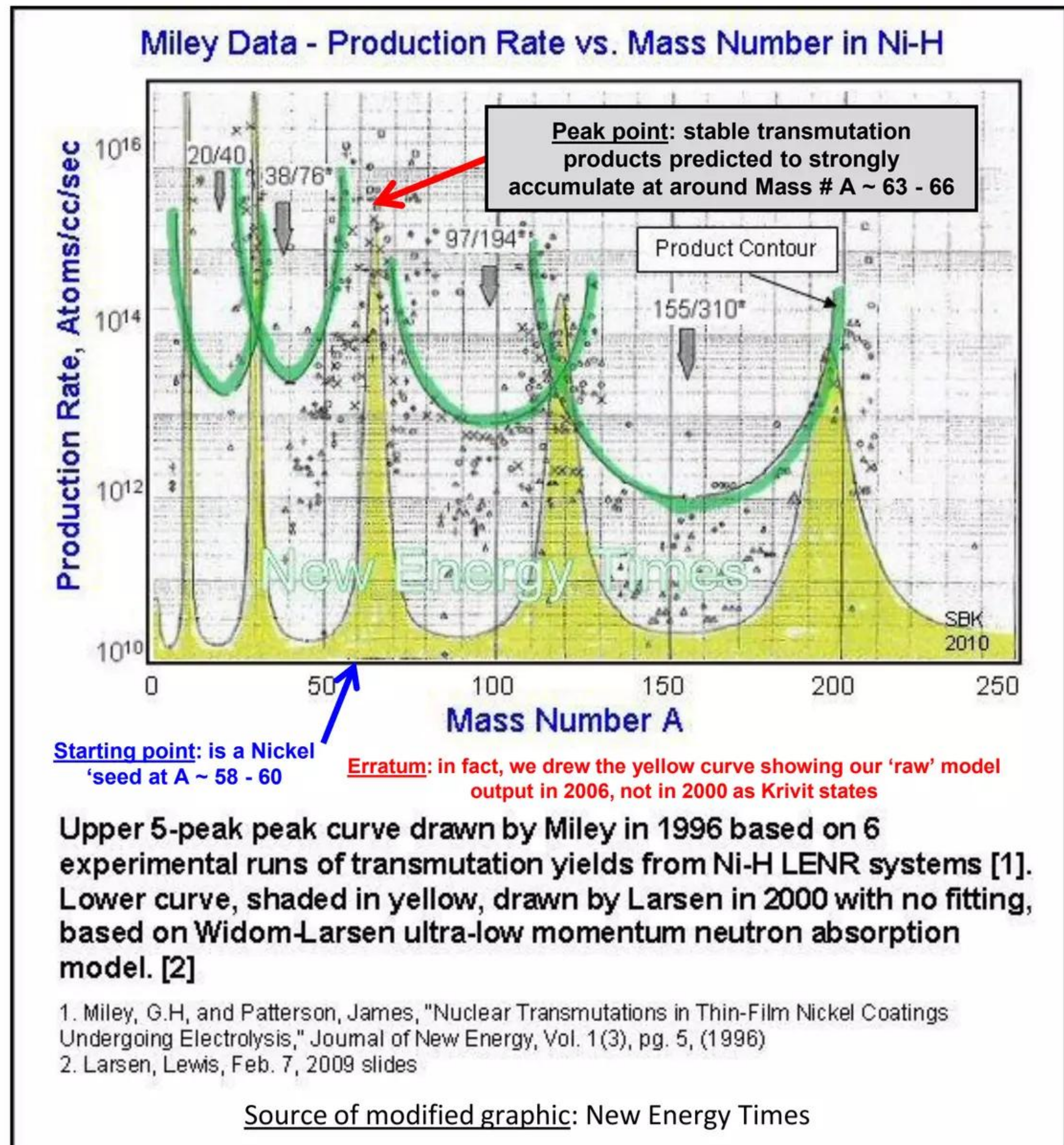


Five-peak mass-spectrum is a 'fingerprint' of ULM neutrons

W-L ULM neutron absorption model (2006) strongly predicts a stable product isotopic abundance peak at around mass # $A \sim 63 - 66$ (i.e., stable isotopes will tend to accumulate around that value of mass), that is, at around $\sim \text{Cu}$ thru $\sim \text{Zn}$, which is clearly observed in Miley's 1996 experimental data shown to the right. The next major mass-peak number where somewhat larger quantities of stable LENR transmutation products are predicted by W-L to accumulate lies out at $A \sim 120$, which is also observed in Miley's (and Mizuno's) data.

Please recall that Miley's ending transmutation product rate data came from multiple P-F type electrolytic Ni-H experimental cell systems that were run for up to several weeks. Also, the beginning Nickel 'seed' in Miley's experiments was a prosaic Ni cathode comprised of (this is starting point):

Ni-58 $\sim 68.0\%$
 Ni-60 $\sim 26.2\%$
 Ni-61 $\sim 1.14\%$
 Ni-62 $\sim 3.63\%$
 Ni-64 $\sim 0.92\%$

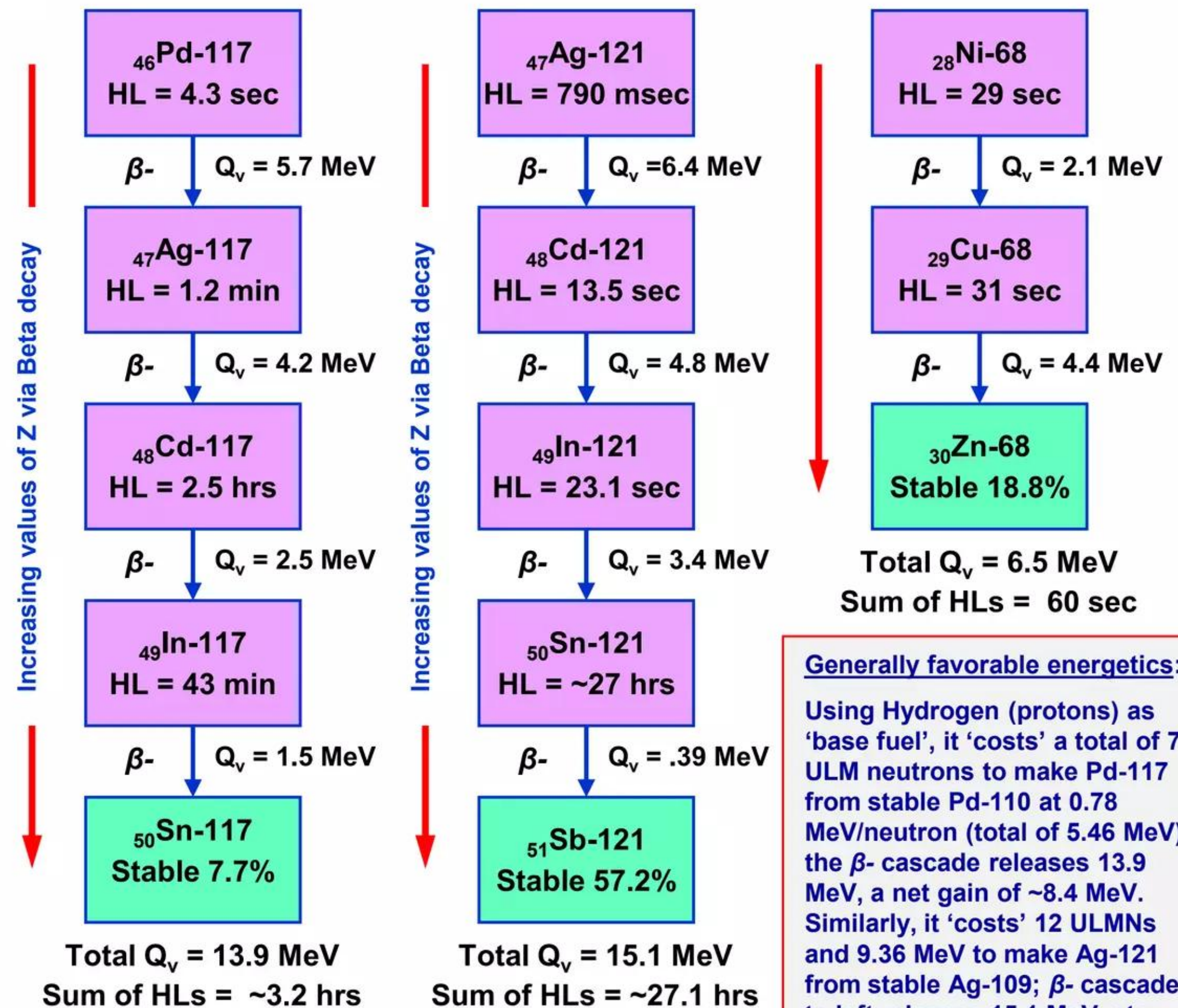


β^- decay chains: very neutron-rich nuclei decay into stable isotopes

Cascades of rapid, energetic β^- -decays are a unique and characteristic feature of LENR systems

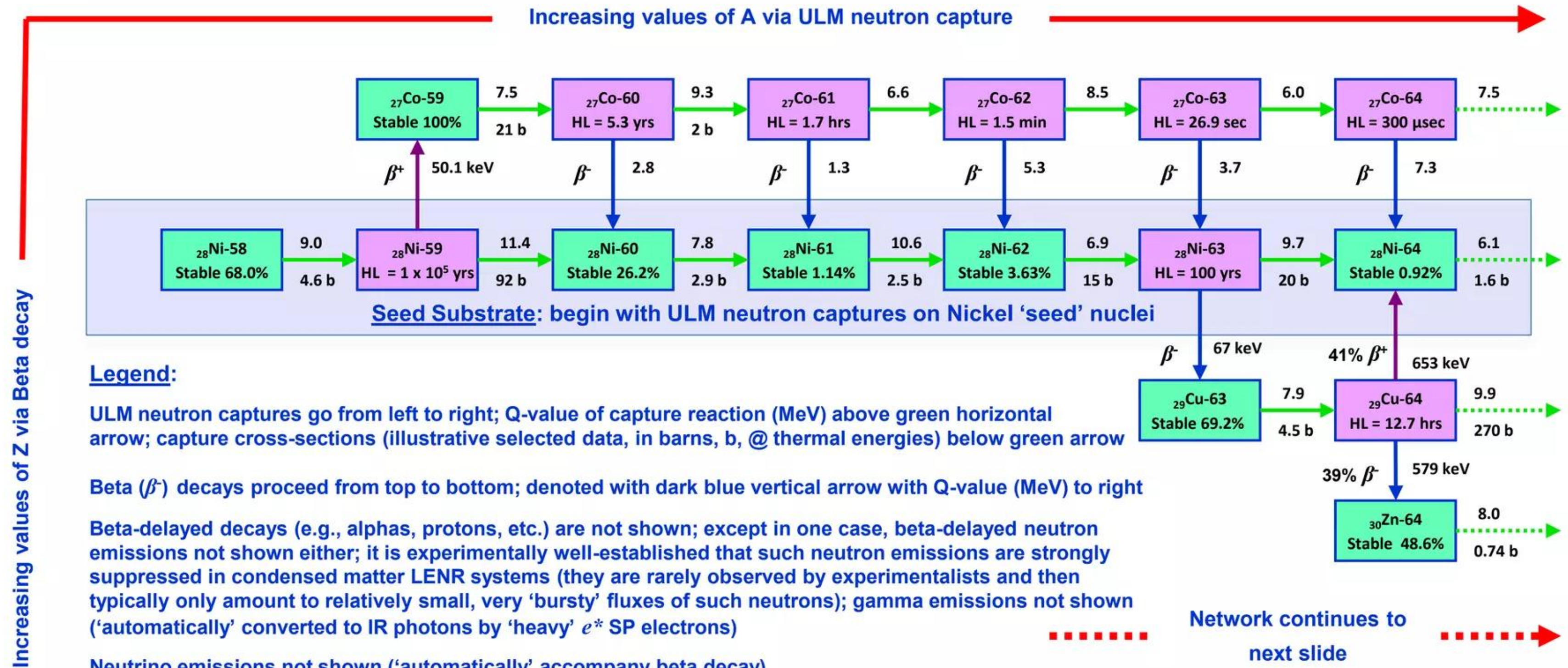
Representative examples of rapid β^- decay cascades:

- ✓ Acc. to W-L theory, over time, large fluxes of ULM neutrons will result in a build-up of large populations of unstable, *very neutron-rich isotopes*
- ✓ At some point, all such n-rich isotopes must decay, often by series of very fast β^- cascades
- ✓ β^- decays release energetic β particles (electrons) that transfer kinetic energy to local matter, heating it up
- ✓ Depending on half-lives, β^- chains can rapidly traverse rows of the periodic table, terminating in production of stable, higher-Z elements. Long-running experiments with large ULMN fluxes may produce a great variety of different elements/isotopes



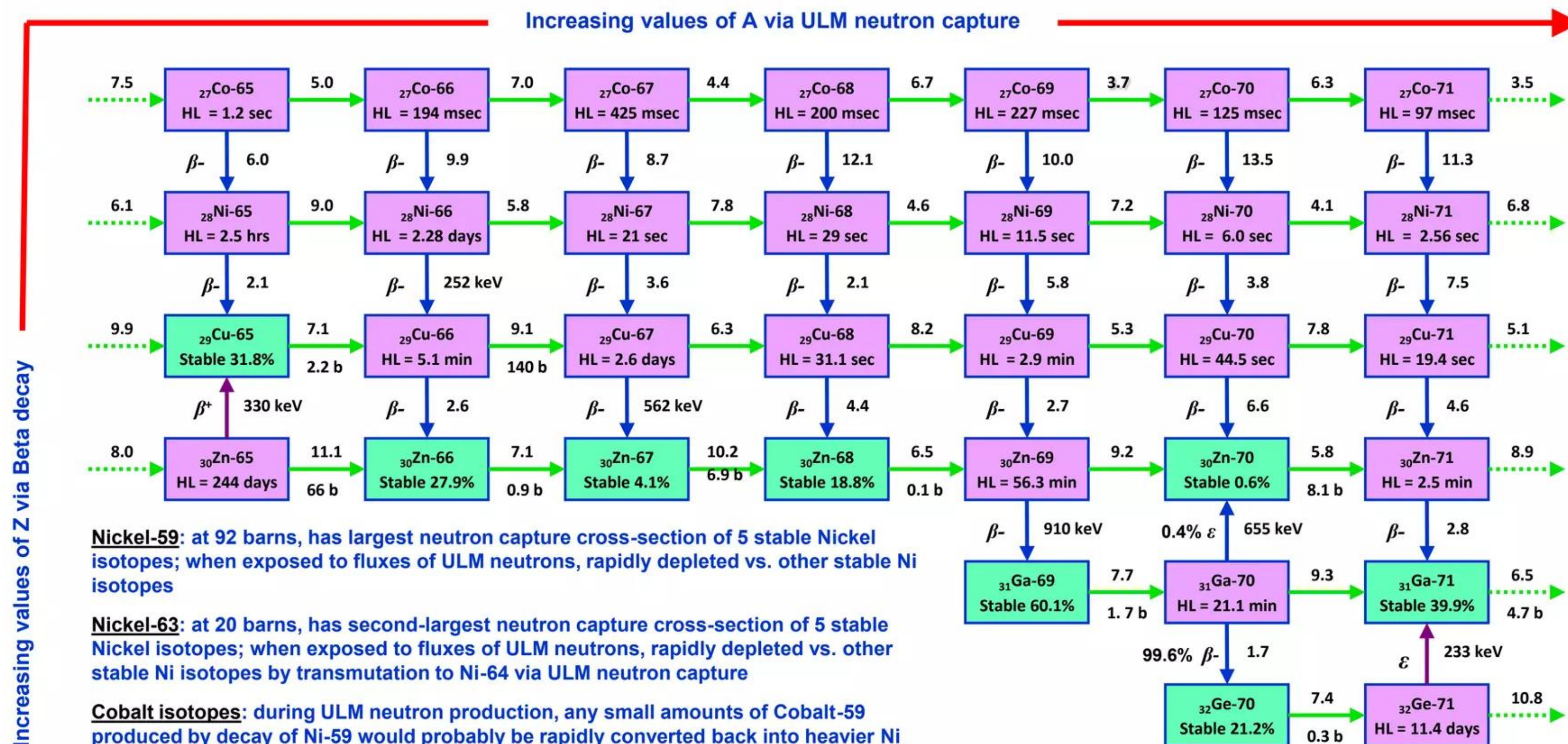
LENR W-L ULM neutron capture on Ni 'seeds,' neutron-rich isotope production, and decays

Stable 'target' seed nuclei on or very near LENR nuclear-active Nickel surfaces: Ni-58, Ni-60, Ni-61, Ni-62, and Ni-64



N.B: in some cases, *Q-values for ULM neutron capture reactions are significantly larger than Q-values for 'competing' beta decay reactions. Also, neutron capture processes are much, much faster (~picoseconds) than most beta decays; if ULM neutron fluxes (rates) are high enough, neutron-rich isotopes of a given element can build-up (move along same row to right on the above chart) >>> faster than beta decays can transmute them to different chemical elements (e.g., move downward to other rows on chart)*

LENR W-L ULM neutron capture on Ni 'seeds,' neutron-rich isotope production, and decays



Nickel-59: at 92 barns, has largest neutron capture cross-section of 5 stable Nickel isotopes; when exposed to fluxes of ULM neutrons, rapidly depleted vs. other stable Ni isotopes

Nickel-63: at 20 barns, has second-largest neutron capture cross-section of 5 stable Nickel isotopes; when exposed to fluxes of ULM neutrons, rapidly depleted vs. other stable Ni isotopes by transmutation to Ni-64 via ULM neutron capture

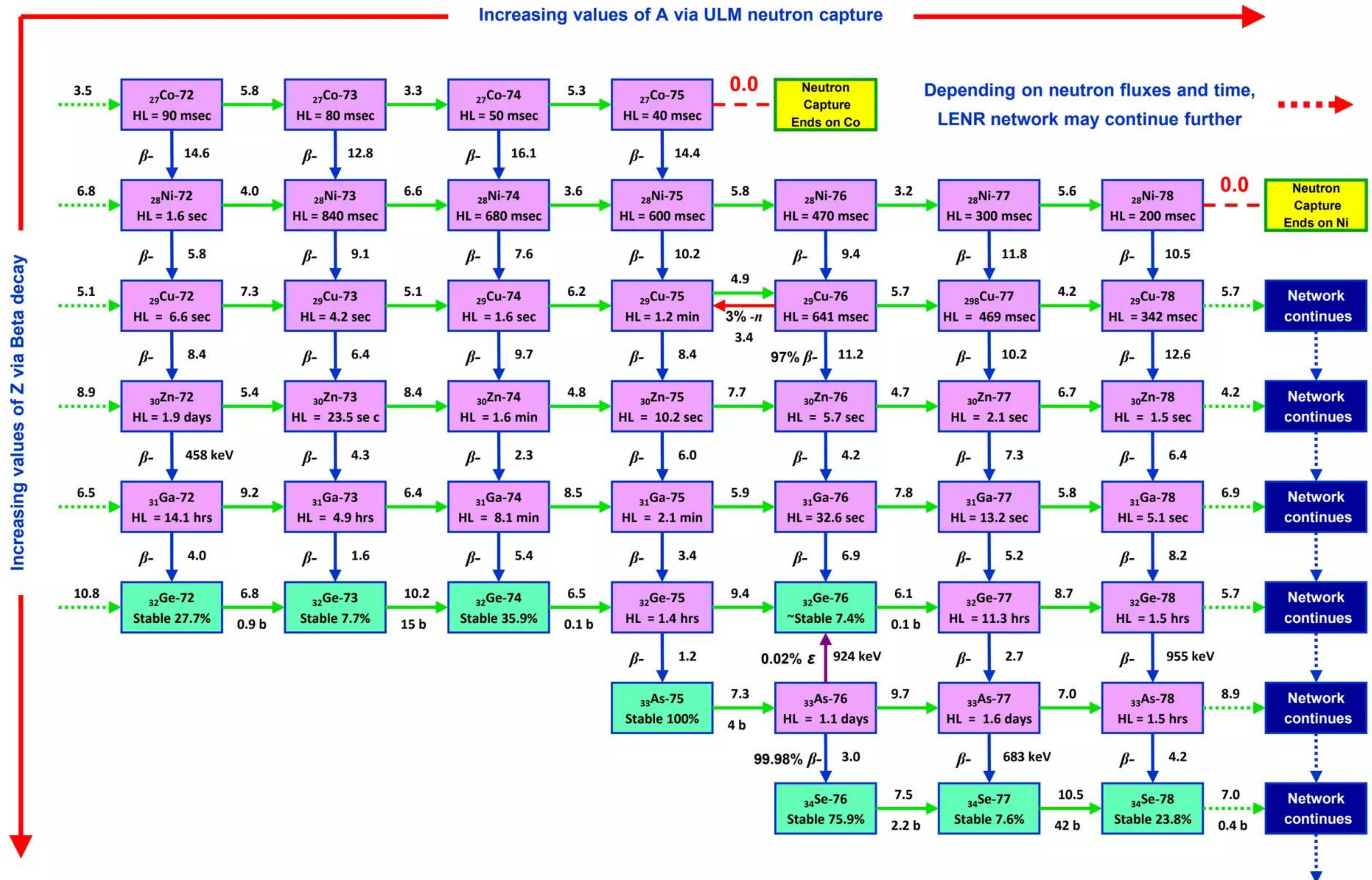
Cobalt isotopes: during ULM neutron production, any small amounts of Cobalt-59 produced by decay of Ni-59 would probably be rapidly converted back into heavier Ni isotopes via ULM neutron capture; prior to Cobalt-60 being converted to Co-61 by neutron capture, any emitted 1.33 and 1.17 MeV Co-60 gammas would be efficiently converted to IR by 'heavy' SP electrons. *To my knowledge, Co-60 gamma signatures have never been claimed to have been observed in any LENR experiment*

Predicted accumulation of stable isotopes at masses A~ 63 - 66: a narrow-width stable isotope abundance peak at around these masses is predicted by the W-L optical model (see previous Slides); model also implies that production of stable isotopes should fall-off relatively rapidly just beyond Copper and Zinc, assuming all other things being equal

Values of A and Z reached by Ni-seed nucleosynthetic networks in experiments depend on overall rates of ULM neutron production (which in turn depend upon the nature of the input energy) and time: the summation across all local neutron fluxes in micron-scale surface regions and their total duration in time (neutron dose history per unit of area) will determine exactly how far a given experiment can proceed into the labyrinth of our above-described Ni-seed network. In gas-phase LENR experiments in which energy is inputted into SP electrons solely through a combination of pressure and temperature, local ULMN fluxes would likely be modest (in comparison to experiments with additional high-current electrical input) and it might likely be difficult for the network to go very far beyond Zinc during a relatively short period of time (say, a few weeks). At the other extreme, electrolytic LENR experiments can sometimes trigger large bursts of ULM neutron production: in one such result, Mizuno went from K to Fe in 2 minutes (Lattice June 25, 2009, SlideShare)

..... Network continues to next slide

LENR W-L ULM neutron capture on Ni 'seeds,' neutron-rich isotope production, and decays



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Conclusions & Intriguing Opportunities for Further Study

- ✓ In condensed matter, W-L production of ULM neutrons creates local, micron-scale regions containing populations of unstable, very neutron-rich nuclei/isotopes; local continuum has densely occupied fermionic states (e.g., SP electrons and neutrons); in parallel, W-L gamma suppression mechanism converts gammas to IR photons with a variable 'tail' in soft X-rays
- ✓ Unlike nuclei lying close to the 'valley of stability,' neutron-rich isotopes have a vastly larger repertoire of different decay and deexcitation channels/modes that can vary quite rapidly and dynamically in time, depending upon the 'sensed' quantum-mechanical 'local continuum environment'; importantly, the closer such unstable nuclei get to the so-called neutron 'dripline,' the more complex and richly varied their behavior can potentially become
- ✓ Think of a neutron-rich nucleus as a collective, many-body quantum system that 'senses' its local continuum through decay channels that are quantum mechanically connected (via entanglement) to the 'outside' world; in effect, such excited nuclei perform dynamic quantum computation in order to 'decide' how to move toward the lowest possible energy state in the least amount of time within given extant constraints (e.g., fermionic decay 'frustration')
- ✓ While many different decay/deexcitation modes are available to neutron-rich nuclei, the most common last stage in their varying and often circuitous 'descent' from excited states to lower, more stable energy states (while maximizing entropy production along the way) are cascades of very fast, spontaneous beta decay chains that eventually terminate in stable isotopes/elements
- ✓ Altogether, this explains why LENR systems emit very little 'hard' neutron or gamma photon radiation and are strongly prone to producing ending arrays of stable isotopes/elements

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Selected Technical Publications

“Ultra Low Momentum Neutron Catalyzed Nuclear Reactions on Metallic Hydride Surfaces”

Eur. Phys. J. C **46**, pp. 107 (March 2006) Widom and Larsen – initially placed on arXiv in May 2005 at http://arxiv.org/PS_cache/cond-mat/pdf/0505/0505026v1.pdf; a copy of the final EPJC article can be found at: <http://www.newenergytimes.com/v2/library/2006/2006Widom-UltraLowMomentumNeutronCatalyzed.pdf>

“Absorption of Nuclear Gamma Radiation by Heavy Electrons on Metallic Hydride Surfaces”

http://arxiv.org/PS_cache/cond-mat/pdf/0509/0509269v1.pdf (Sept 2005) Widom and Larsen

“Nuclear Abundances in Metallic Hydride Electrodes of Electrolytic Chemical Cells”

http://arxiv.org/PS_cache/cond-mat/pdf/0602/0602472v1.pdf (Feb 2006) Widom and Larsen

“Theoretical Standard Model Rates of Proton to Neutron Conversions Near Metallic Hydride Surfaces”

http://arxiv.org/PS_cache/nucl-th/pdf/0608/0608059v2.pdf (v2. Sep 2007) Widom and Larsen

“Energetic Electrons and Nuclear Transmutations in Exploding Wires”

http://arxiv.org/PS_cache/arxiv/pdf/0709/0709.1222v1.pdf (Sept 2007) Widom, Srivastava, and Larsen

“Errors in the Quantum Electrodynamic Mass Analysis of Hagelstein and Chaudhary”

http://arxiv.org/PS_cache/arxiv/pdf/0802/0802.0466v2.pdf (Feb 2008) Widom, Srivastava, and Larsen

“High Energy Particles in the Solar Corona” http://arxiv.org/PS_cache/arxiv/pdf/0804/0804.2647v1.pdf

(April 2008) Widom, Srivastava, and Larsen

“A Primer for Electro-Weak Induced Low Energy Nuclear Reactions” Srivastava, Widom, and Larsen

Pramana – Journal of Physics **75** pp. 617 (October 2010) <http://www.ias.ac.in/pramana/v75/p617/fulltext.pdf>

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“No single solution will defuse more of the Energy-Climate Era’s problems at once than the invention of a source of abundant, clean, reliable, and cheap electrons. Give me abundant clean, reliable, and cheap electrons, and I will give you a world that can continue to grow without triggering unmanageable climate change. Give me abundant clean, reliable, and cheap electrons, and I will give you water in the desert from a deep generator-powered well. Give me abundant clean, reliable, and cheap electrons, and I will put every petrodictator out of business. Give me abundant clean, reliable, and cheap electrons, and I will end deforestation from communities desperate for fuel and I will eliminate any reason to drill in Mother Nature’s environmental cathedrals. Give me abundant clean, reliable, and cheap electrons, and I will enable millions of the earth’s poor to get connected, to refrigerate their medicines, to educate their women, and to light up their nights.”

Thomas Friedman, “Hot, Flat, and Crowded” 2008 pp.186